Substorms are a major cause of large geomagnetically induced currents

1

Ari Viljanen and Eija Tanskanen Finnish Meteorological Institute

Antti Pulkkinen UMBC/GEST at NASA/GSFC

• Substorm:

Negative bay > 100 nT with a clear onset when the IL index rapidly decreases at least 100 nT within 10 min $(IL(t) = \min\{B_x(t)\})$

- Data: 833 substorms in 1997 and 1999
- Two categories: non-storm (Dst>-40 nT) (696 events) storm-time (Dst<-40 nT) (137 events)

• $d\mathbf{H}/dt$ is a good activity indicator of geomagnetically induced currents (GIC)

IMAGE Magnetometer Network 50° 350° NAL LYR 75[°] HOR HOF BJN 🔵 70°. LOZ RVK OUJ LYC C KAR 🧯 UPS

October 2004





X and dX/dt at Nurmijärvi, and GIC along the natural gas pipeline.



Pirttikoski transformer in northern Finland (CGM lat. 63 deg N).



Average eastward equivalent current density after the onset. Dots show the location of the maximum westward current density. There is a rapid intensification and southward transition after the onset.







 \rightarrow completely different ground electric fields and GIC

Note: further complexity due to a nonuniform conductivity of the earth.

Can we predict $d\mathbf{H}/dt$ (or GIC) during substorms?

- Ionosphere has a short memory of $d\mathbf{H}/dt$ (less than 2 minutes)
- Preference for any particular direction or length scale of $d\mathbf{H}/dt$ decreases during substorms
- It may be impossible to predict $d\mathbf{H}/dt$ in a deterministic way

Pulkkinen et al., 2006: Spatiotemporal scaling properties of the ground geomagnetic field variations. JGR, doi:10.1029/2005JA011294.

However, there is some hope to estimate GIC levels even with quite incomplete knowledge of $d\mathbf{H}/dt$:

Pulkkinen et al., 2006: Estimation of geomagnetically induced current levels from different input data. Space Weather, 4, S08005, doi:10.1029/2006SW000229.

Conclusions

- A few minutes after the onset is the most probable time of large $d\mathbf{H}/dt$ and GIC during (storm-time) substorms.
- The probability of large $d\mathbf{H}/dt$ increases with increasing $|\mathbf{H}|$. For a given $|\mathbf{H}|$, there seems to be a fairly well-defined lower limit of $\max(|d\mathbf{H}/dt|)$.
- Large $d\mathbf{H}/dt$ can occur due to two very different spatial structures:
 - intensification of the main electrojet \rightarrow smooth $d\mathbf{H}/dt$ pattern
 - small-scale radiply varying structures \rightarrow complicated $d\mathbf{H}/dt$ pattern

Future work: study all substorms of a complete sunspot cycle (1996–2006).

Reference: Viljanen et al., 2006: Relation between substorm characteristics and rapid temporal variations of the ground magnetic field. Ann. Geophys., SRef-ID: 1432-0576/ag/2006-24-725.